

Review of Lunar Polar Illumination



Ben Bussey

and

Angela Stickle

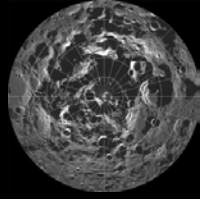
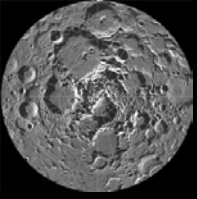
Space Exploration Sector

*The Johns Hopkins University Applied
Physics Laboratory*

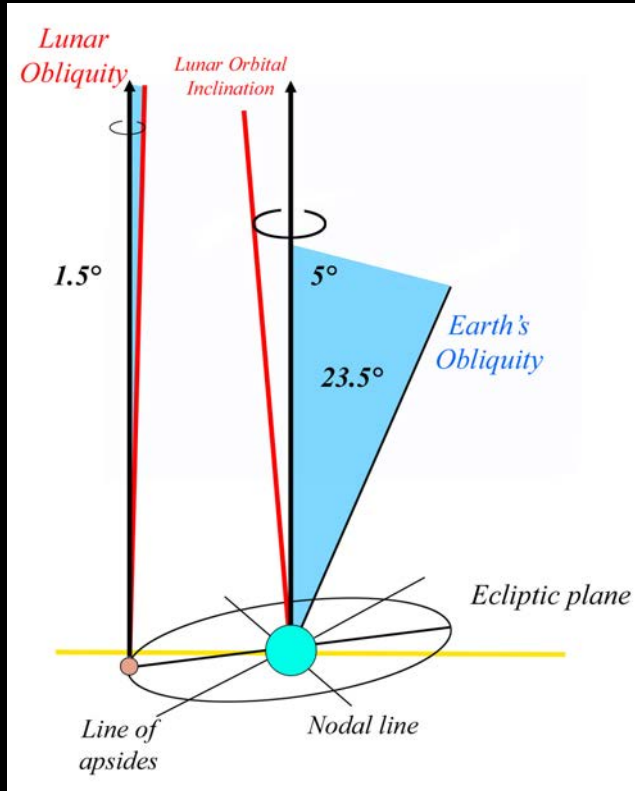
ben.bussey@jhuapl.edu

301-906-1174

angela.stickle@jhuapl.edu

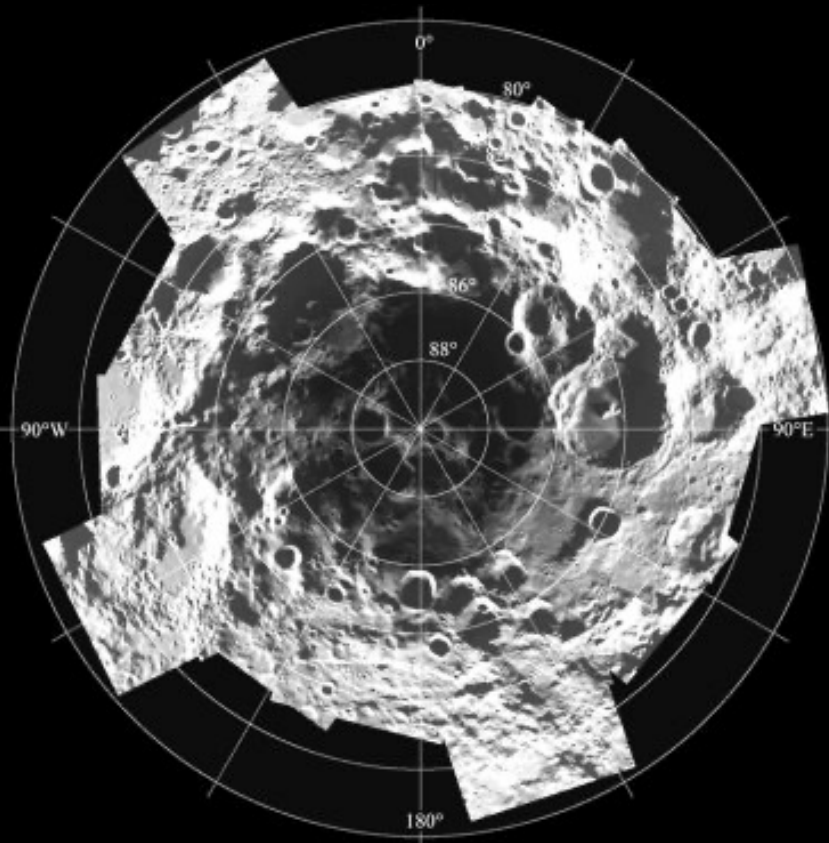


Lunar Poles

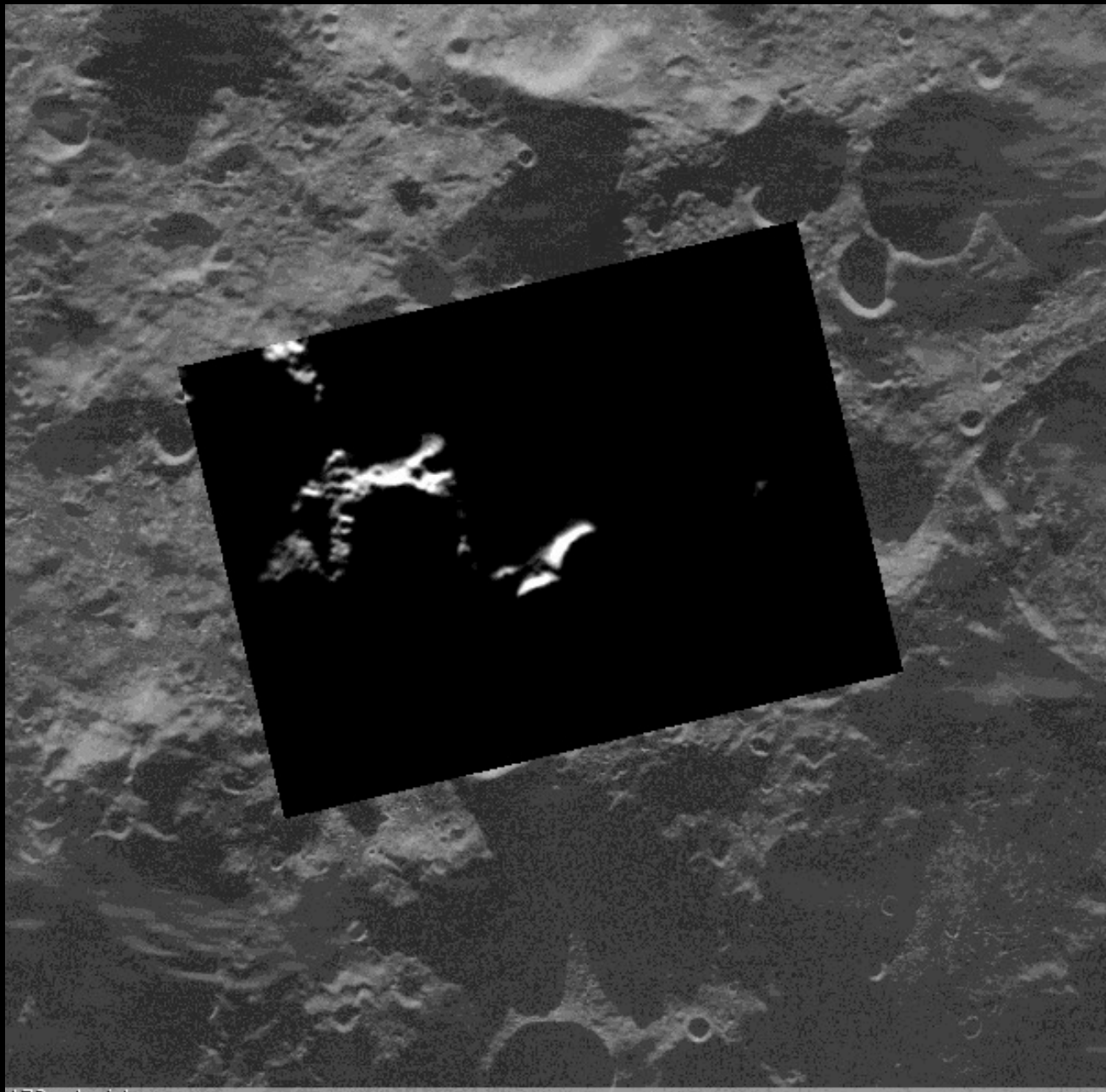


- Spin axis 1.5° from the perpendicular to the ecliptic plane results in special illumination conditions
- Permanently shadowed regions are cold traps ($< 100\text{K}$) and are possible locations of ice deposits
- First mention that permanent shadowed regions could exist was by Urey in his 1952 book *The Planets*

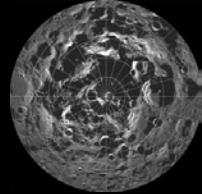
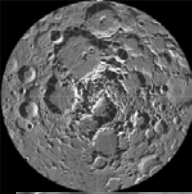
Clementine's View of the Poles



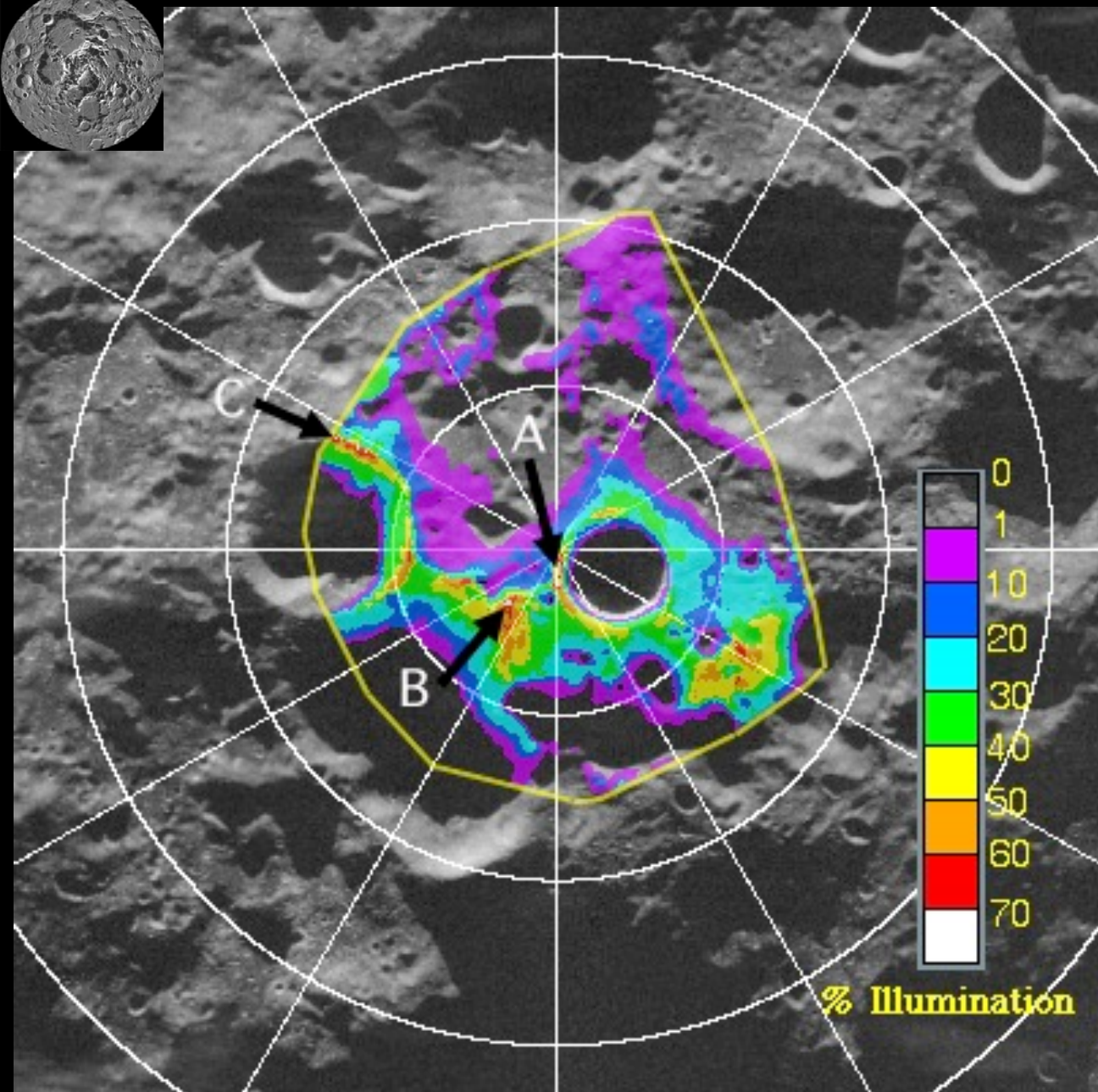
- Imaged each pole every 10 hours for 71 days
- Data collected during winter in the southern hemisphere
- High resolution data collected at 15 m/pixel
- Shoemaker 1994 Science paper estimated at least 30,000 km² of shadow in the south polar region



173.cub.pict



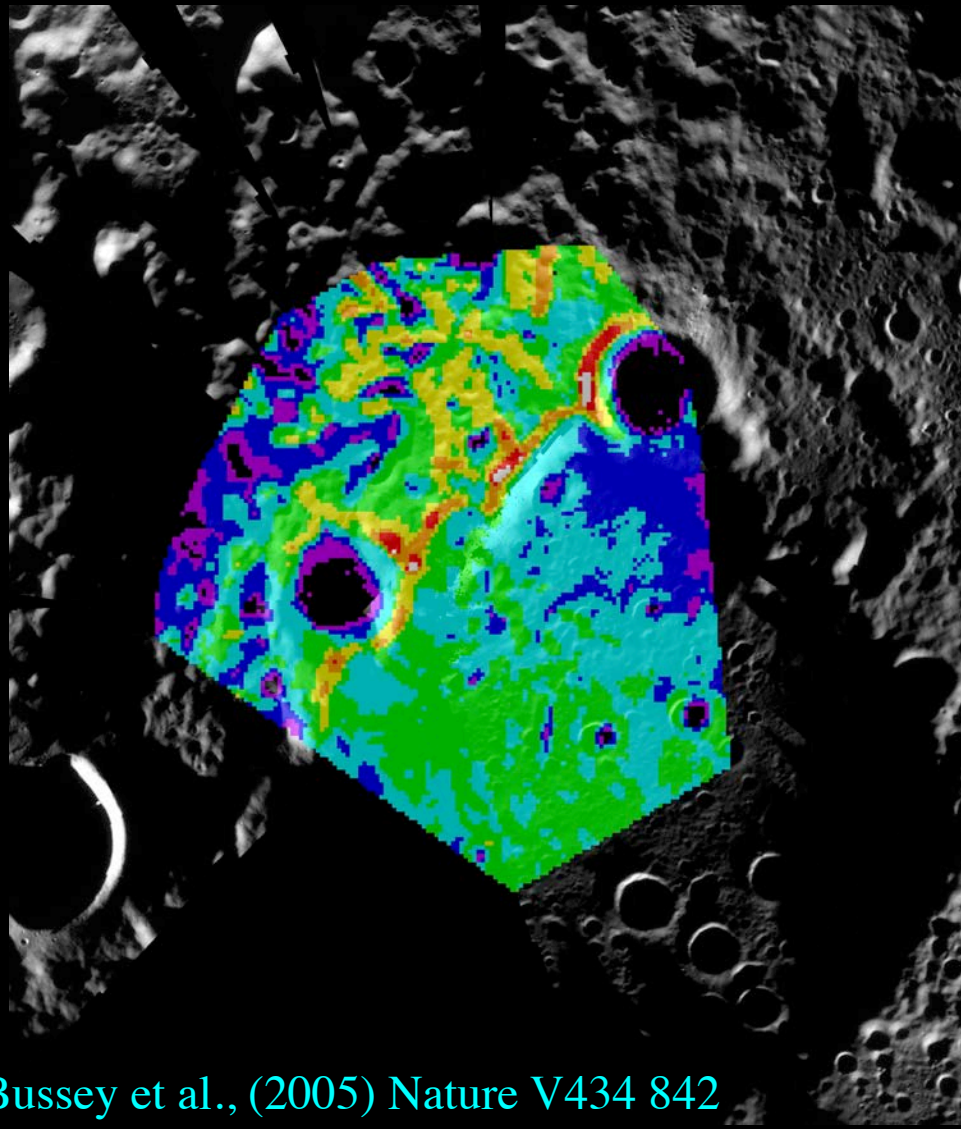
Quantitative Illumination Map



- No constant illumination
- A,B,C lit more than 70% of a winter day
- A&B collectively lit > 98%

Bussey et al., (1999) GRL V26 No. 9, 1187-1190

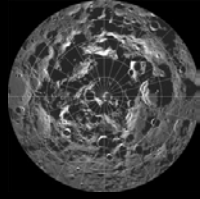
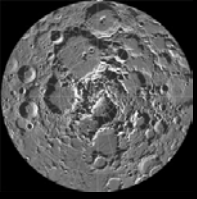
North-Pole Illumination Map



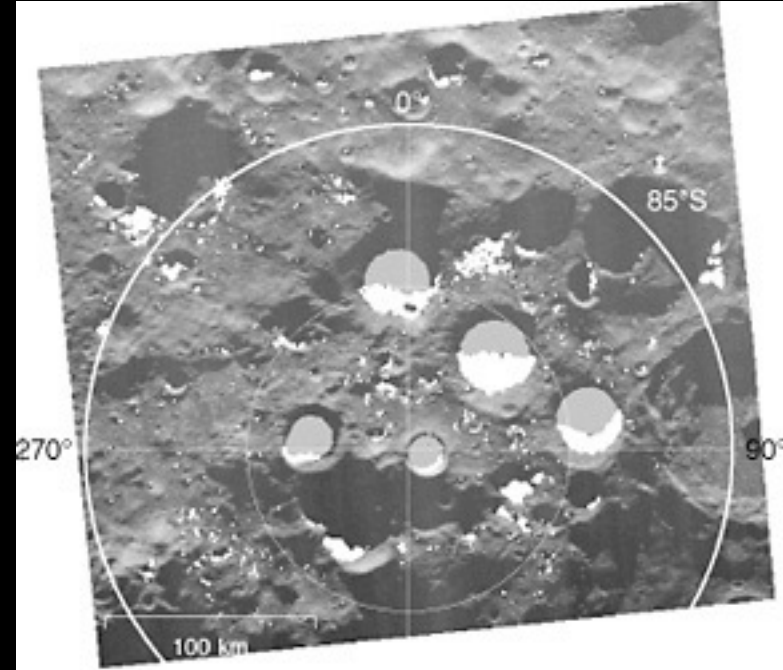
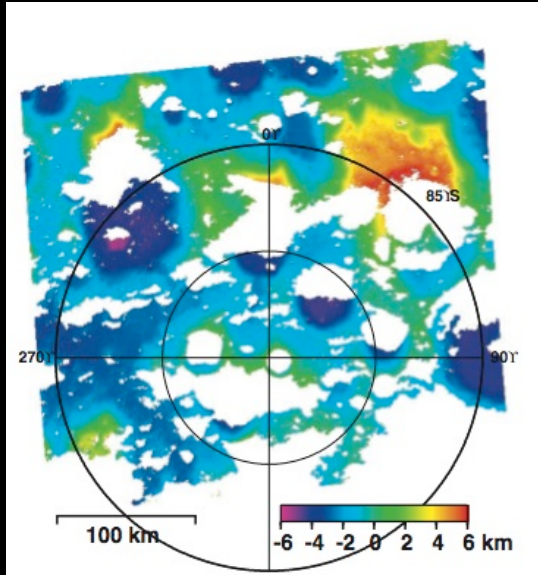
- Four places on the rim of Peary crater were constantly illuminated during a lunar summer day
- All are in close proximity with permanently shadowed regions.



Bussey et al., (2005) Nature V434 842



Earth-Based Radar

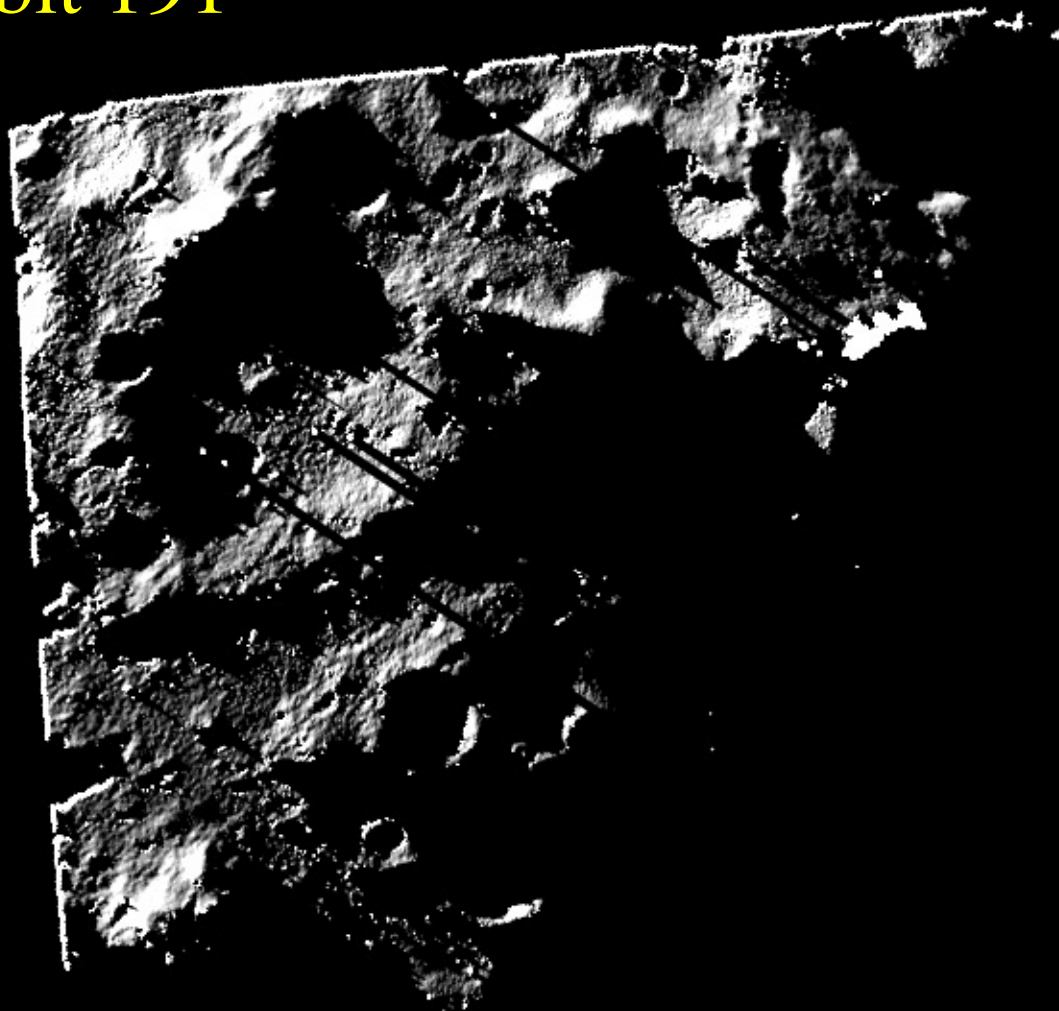


- Made from Goldstone data
- 150 m Spatial, 50 m vertical DEM
- Mapped locations of permanent shadow

Details in Margot et al (1999) Science V 284 1658-1660

Orbit 191

Clementine stereo



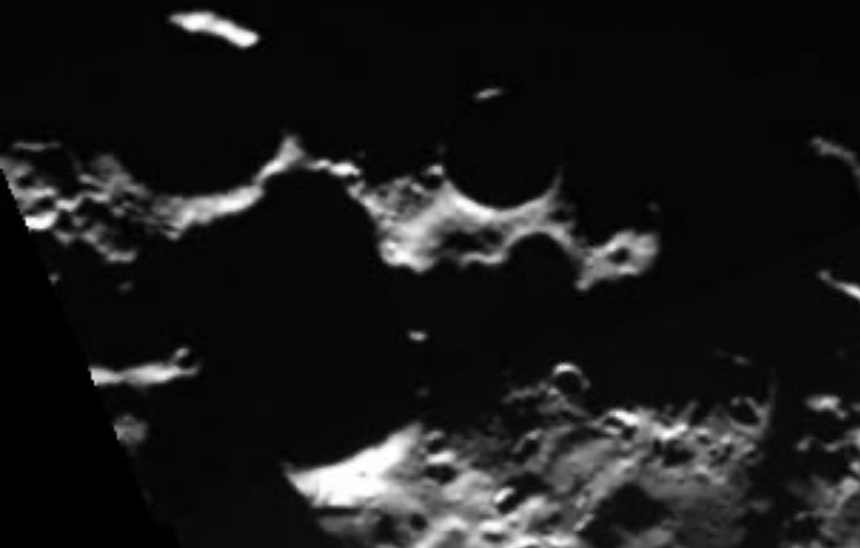
Clementine UVVIS

Radar topography

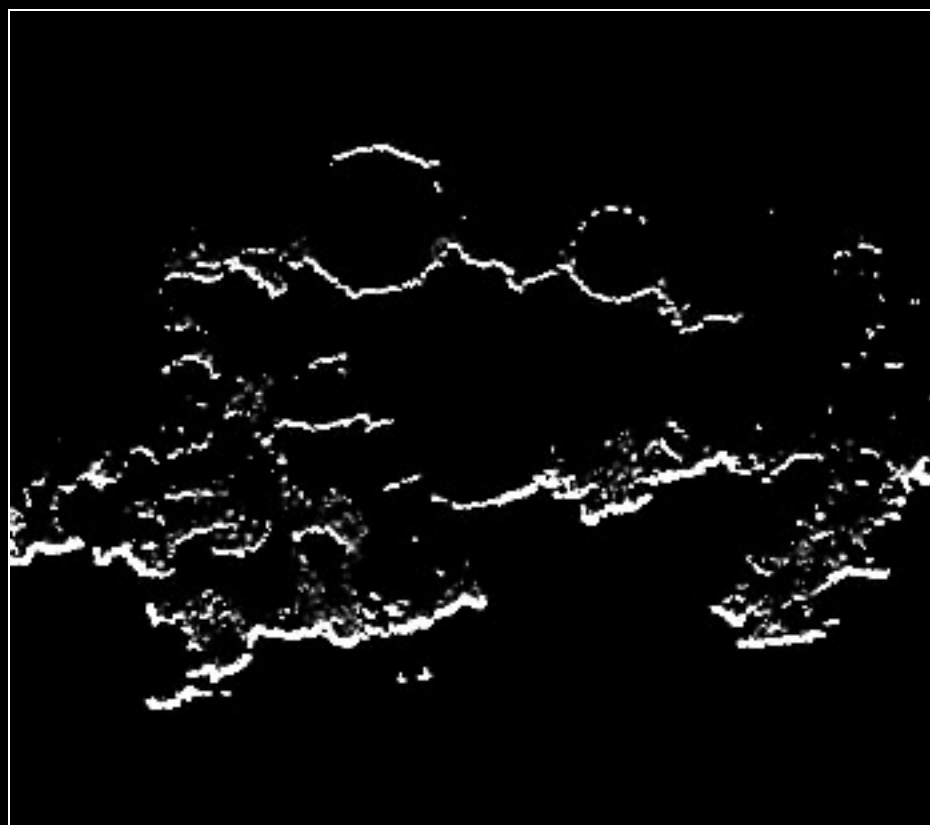
Orbit 243



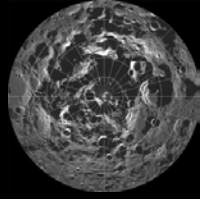
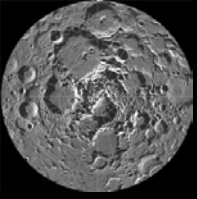
Clementine stereo



Clementine UVVIS

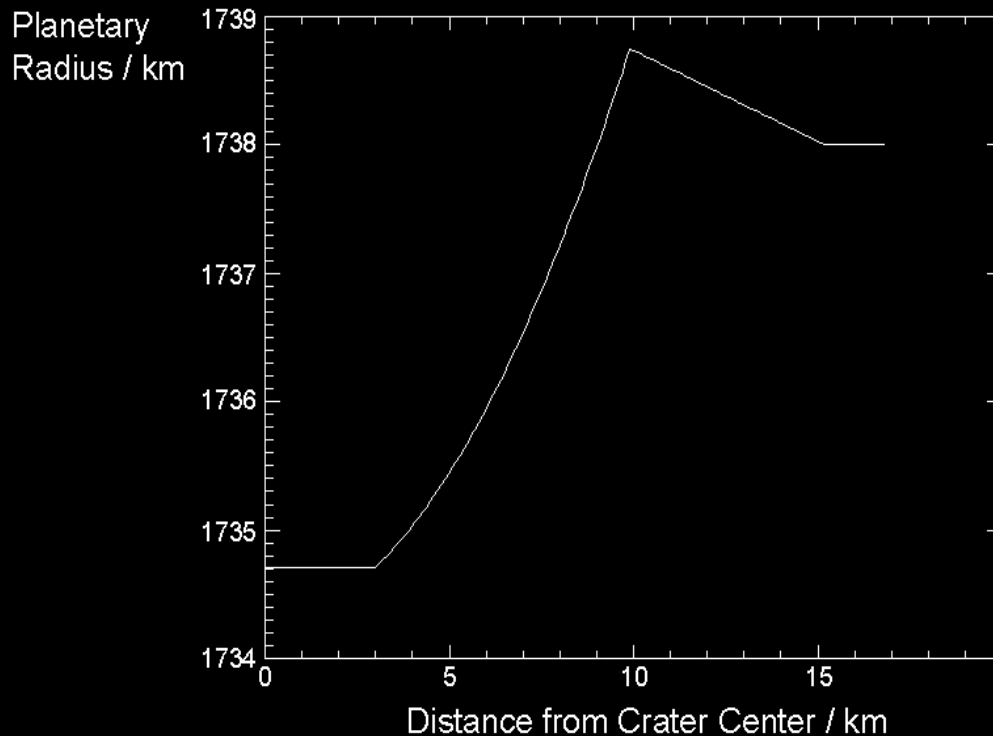


Radar topography



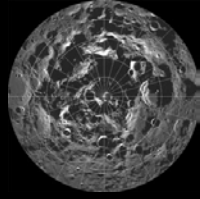
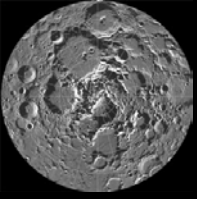
Simulated Simulations

20 km Crater Profile



- Crater profiles for lunar craters are known (e.g. Pike, 1977).
- Investigate amount of permanent shadow as a function of size, latitude and season.

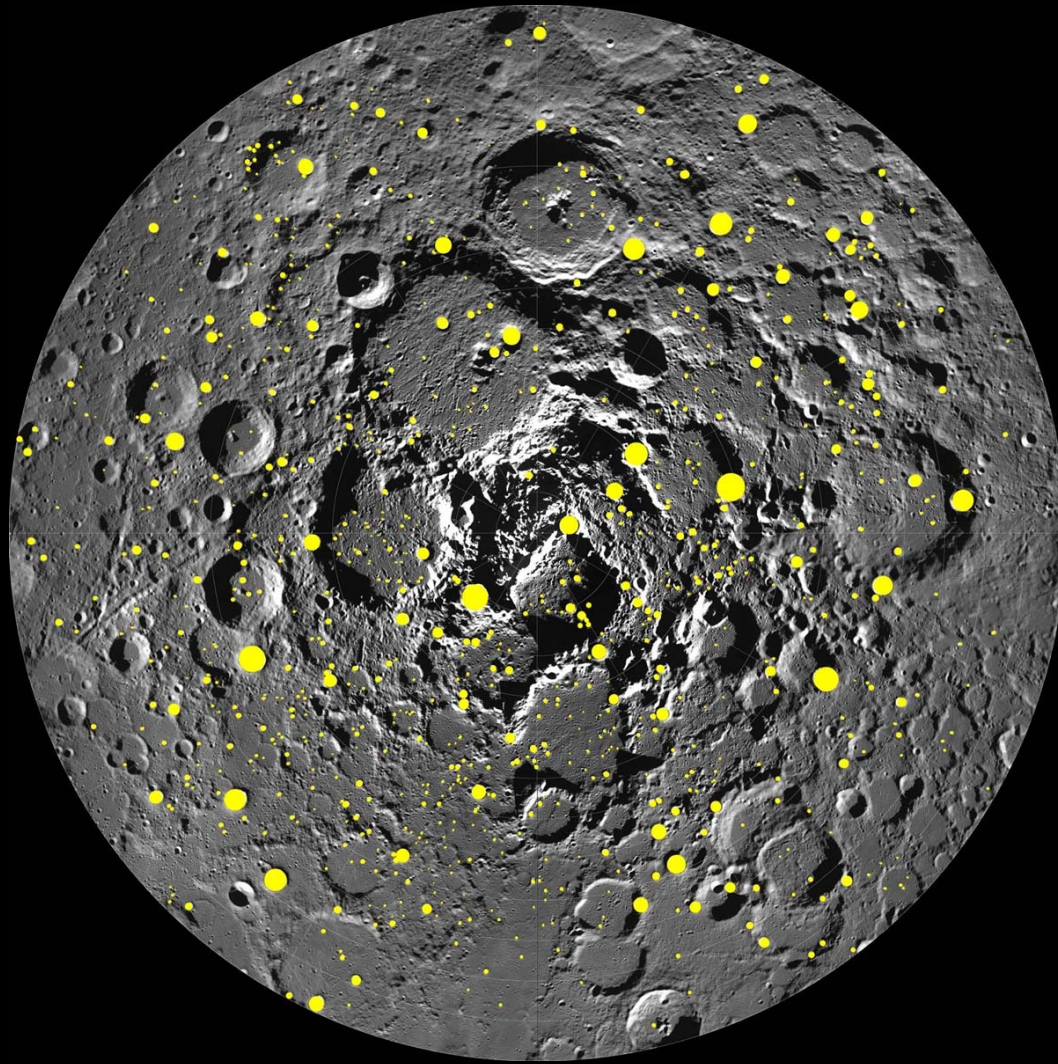
Topography simulations



$$S=(0.9465\times D) + (0.0202\times\theta^2) - (0.009258\times\theta\times D) - 78.06$$

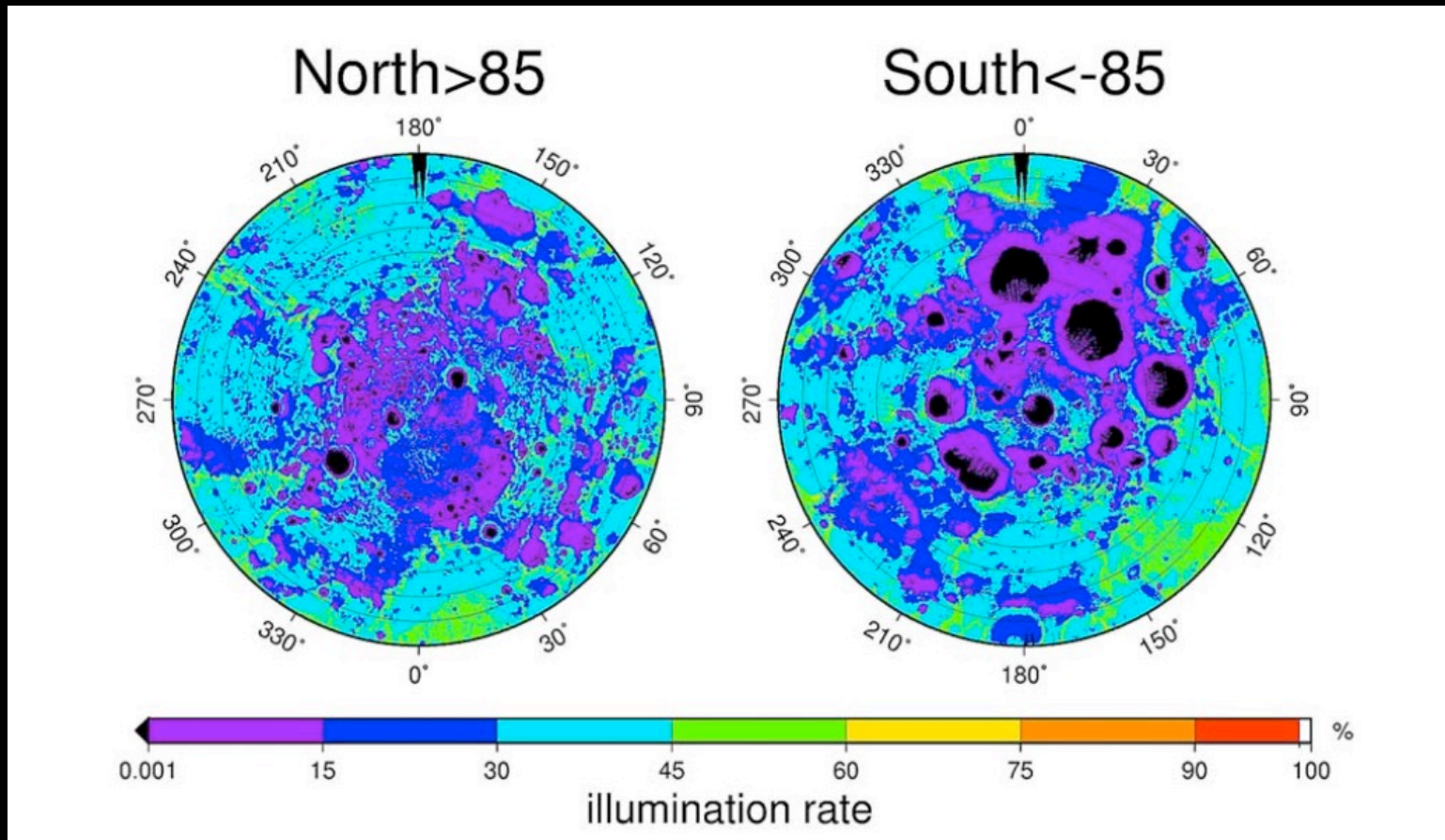
- Craters in the range 2.5 to 20 km were studied
- Craters placed every 1° latitude from 70° to 90°
- A series of runs are conducted to calculate the extent of permanent shadow

Permanent Shadow



- 1000's km² of permanent shadow in simple craters at both poles
- Permanent shadow exists in simple craters at long distances from the pole
- Represent possible locations of volatile deposits

Kaguya



- No 100% illuminated areas
- Both poles have locations that receive illumination $> 80\%$ of the time
- See Noda et al., (2008) GRL V35 L24203

Kaguya-Clementine Comparison

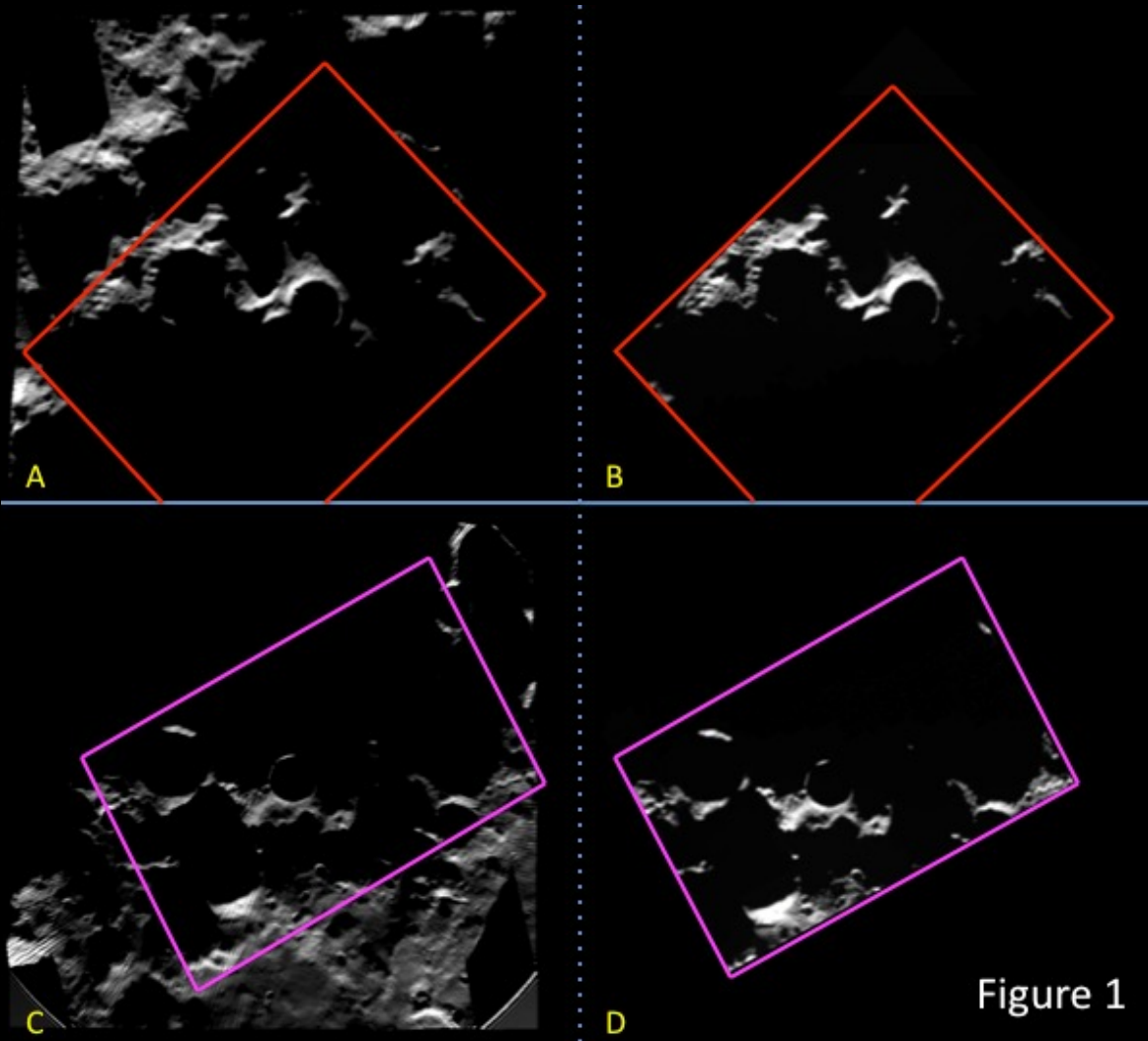
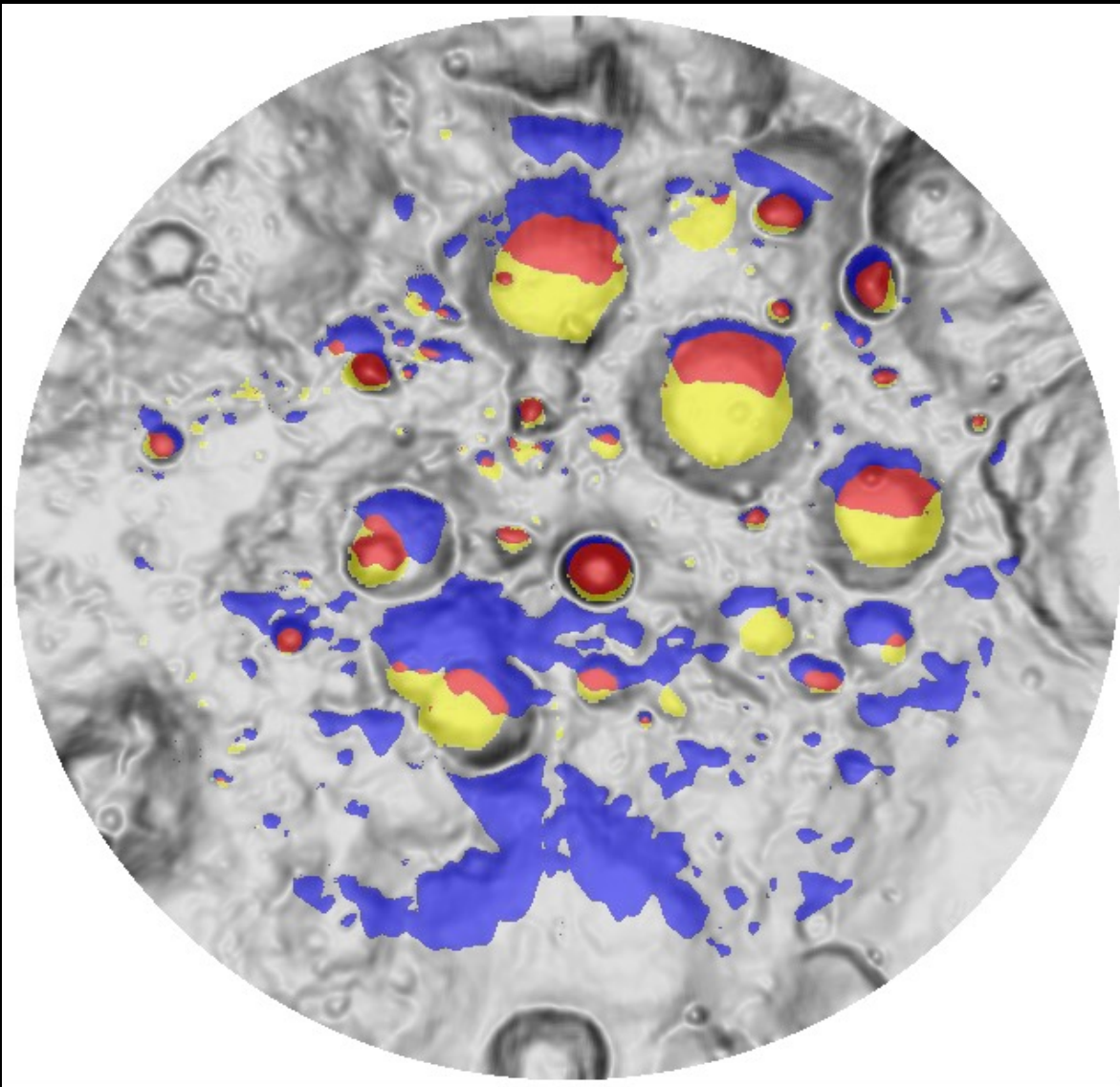


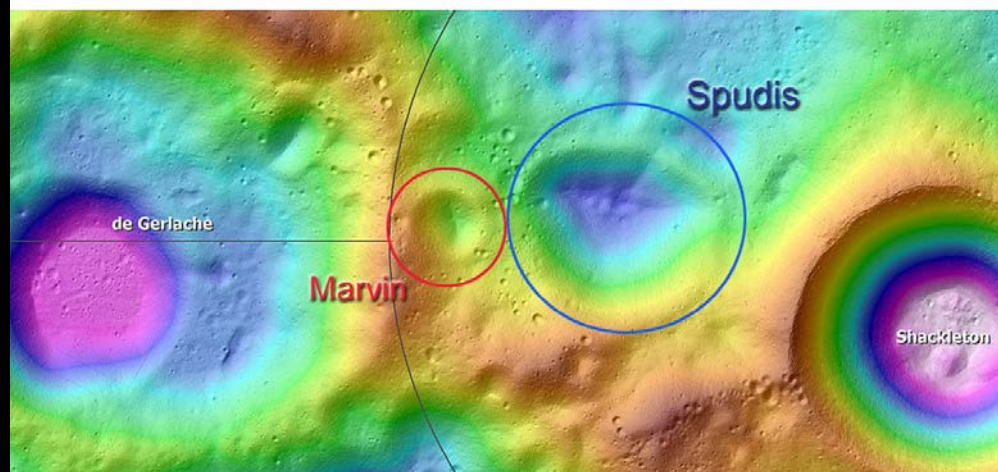
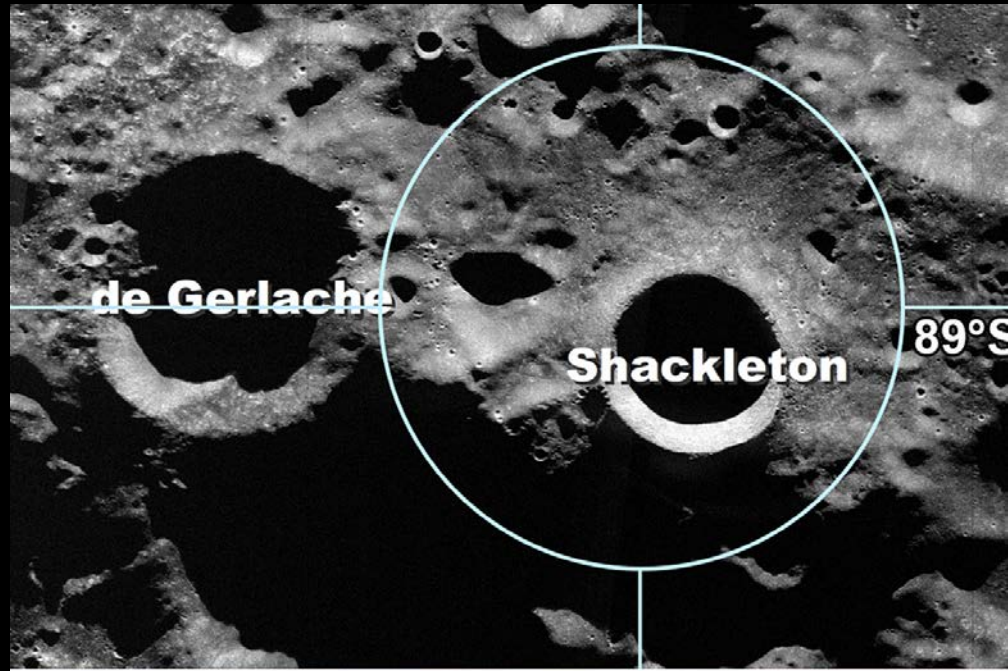
Figure 1

Earth & Sun Shadows



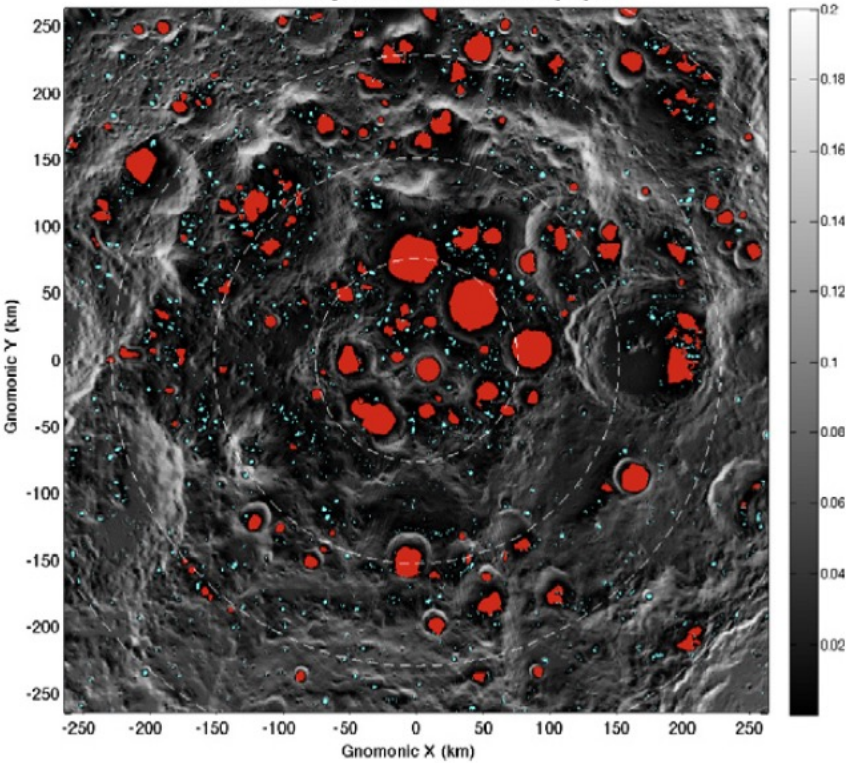
- Earth Shadowed
 - Blue & Red
- Sun Shadowed
 - Yellow & Red

LRO



LOLA

South Pole average incident flux over four 18.6yr cycles

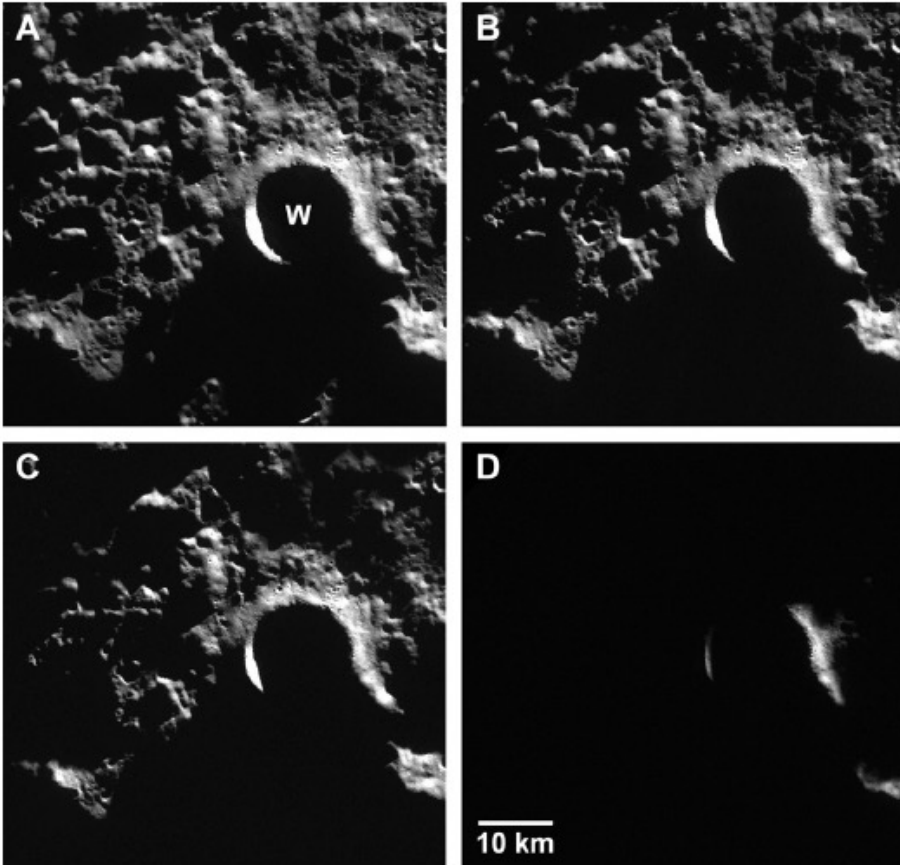


Region	North	South
$\sim 80^\circ$	12866	16055
$> 82.5^\circ$	8881	12202
$> 85^\circ$	4764	7024
$> 87.5^\circ$	1769	3660
$> 87.5^\circ > 1 \text{ km}^2$	1665	3632
$> 87.5^\circ > 2 \text{ km}^2$	1496	3588
$> 87.5^\circ > 5 \text{ km}^2$	1236	3516
$> 87.5^\circ > 10 \text{ km}^2$	983	3397
$> 87.5^\circ > 20 \text{ km}^2$	575	3306
$> 87.5^\circ > 50 \text{ km}^2$	523	3187

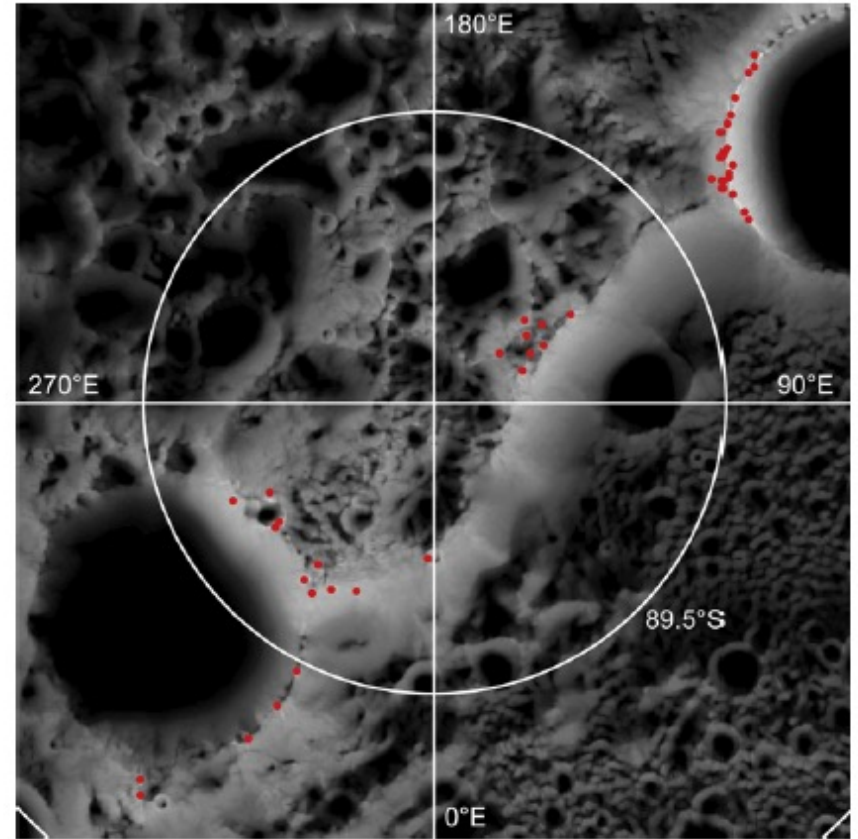
From Mazarico et al., Icarus, 2011

LROC

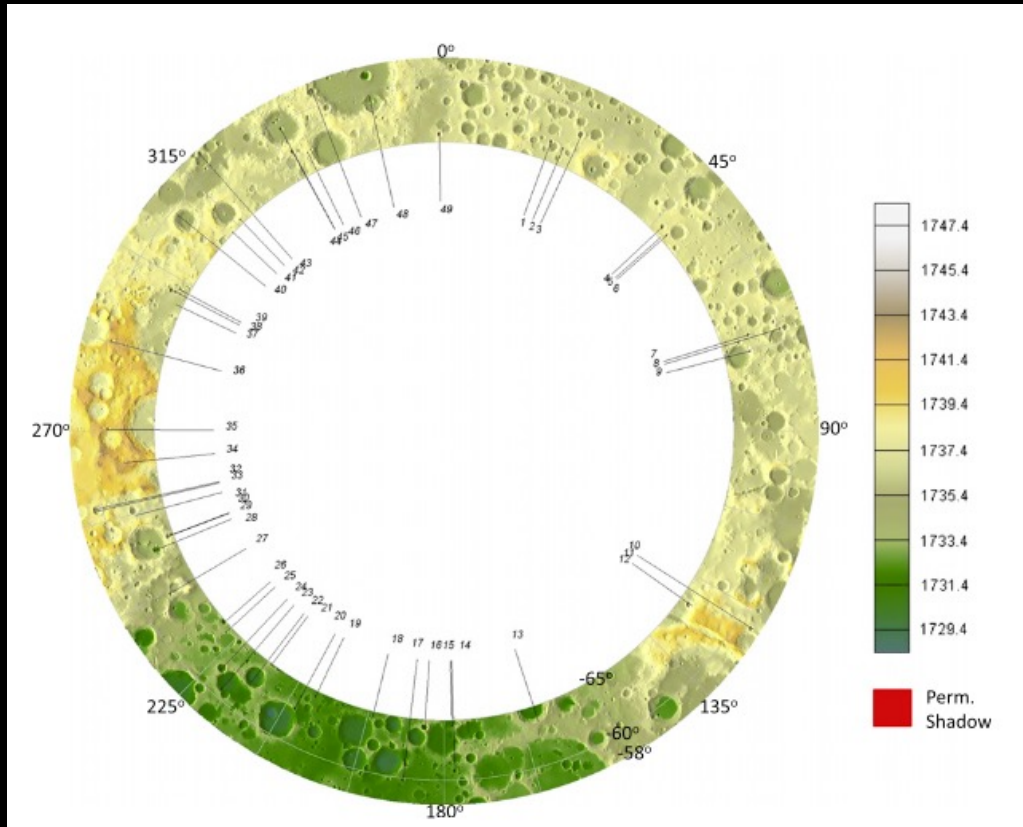
E.J. Speyerer, M.S. Robinson/Icarus 222 (2013) 122–136



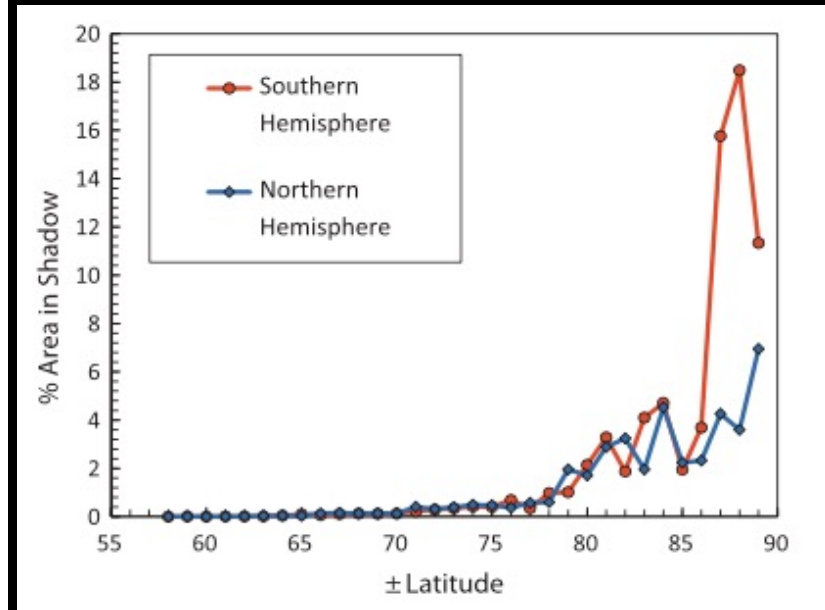
E.J. Speyerer, M.S. Robinson/Icarus 222 (2013) 122–136



LOLA & LROC

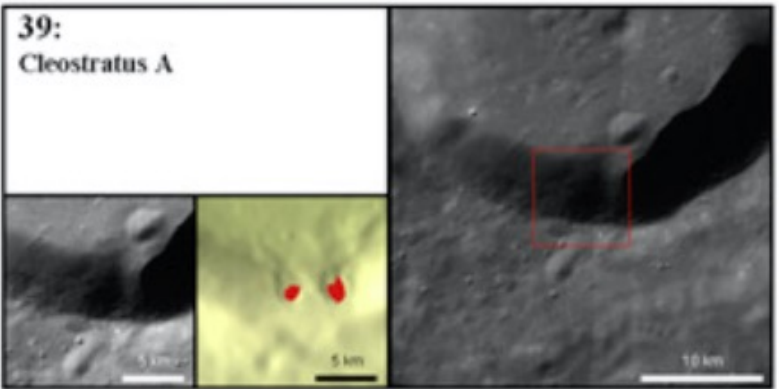


58 - 65° S

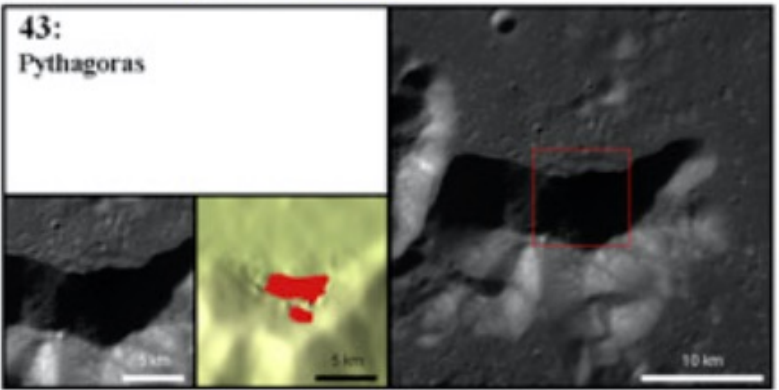


From McGovern et al., Icarus, 2013

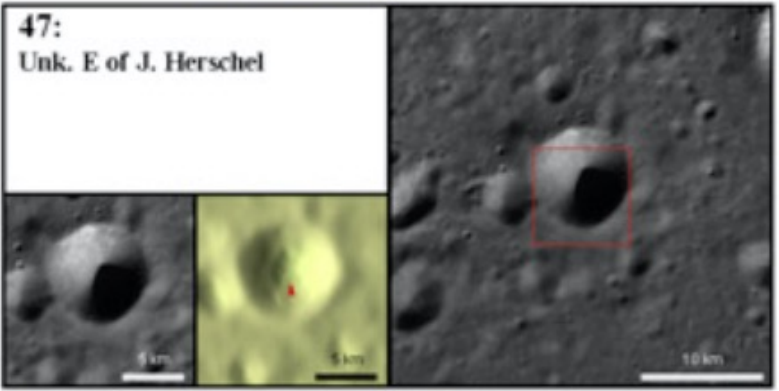
LOLA & LROC



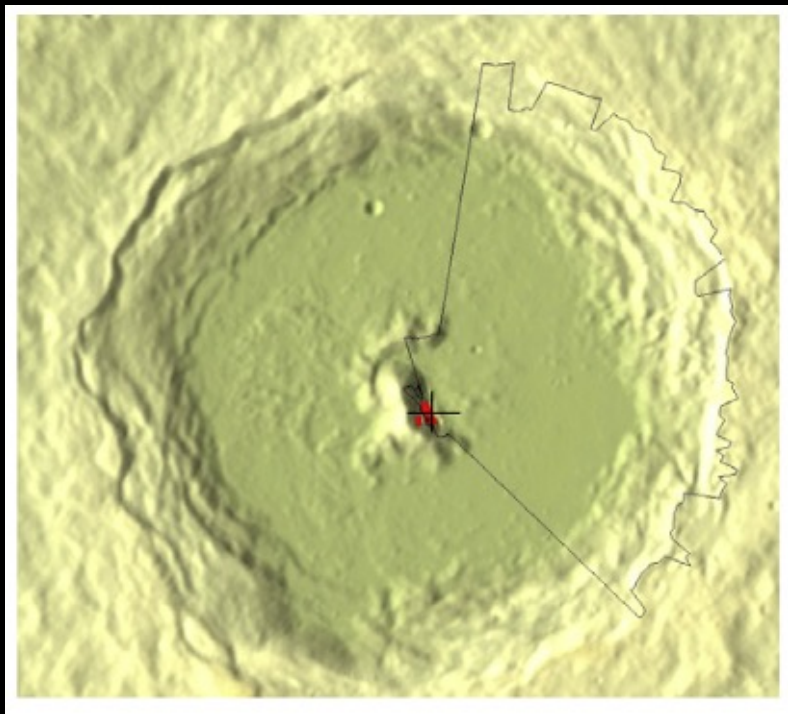
62.2° N



63.5° N



64.9° N

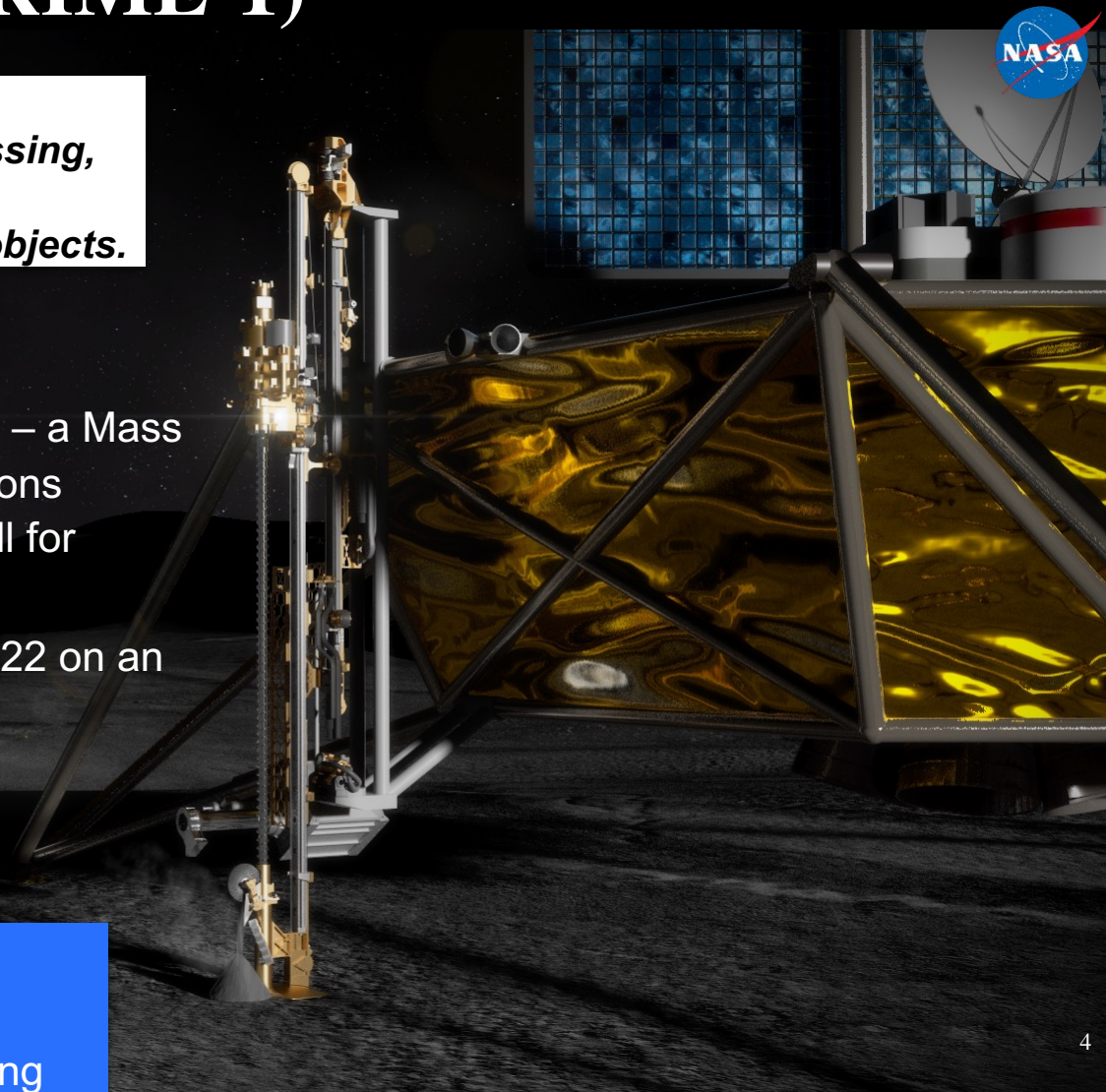


Polar Resources Ice Mining Experiment-1 (PRIME-1)

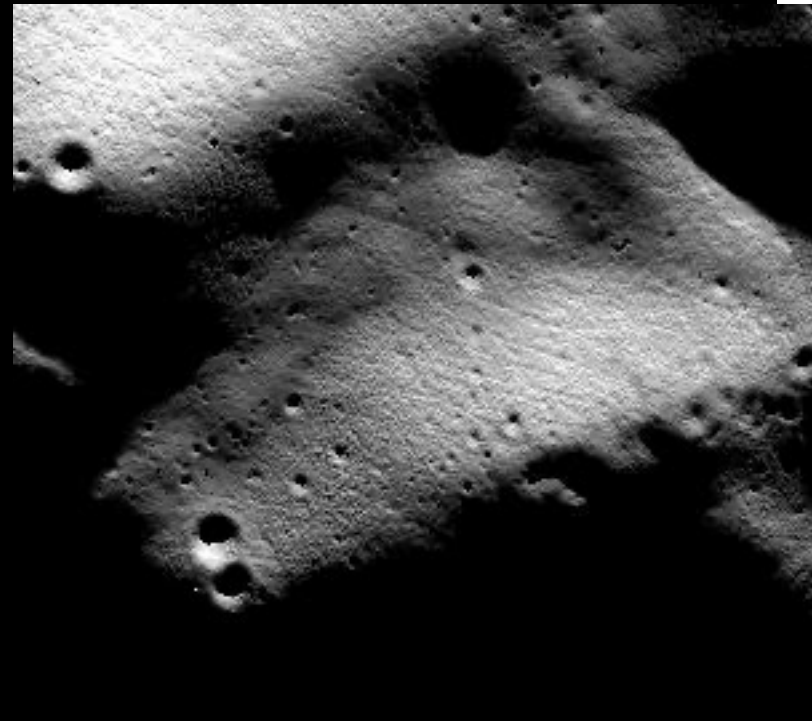
Space Technology is developing technologies for the collection, processing, storing and use of material found or manufactured on other astronomical objects.

- Consists of two high-TRL instruments – a Mass Spectrometer observing lunar operations (MSolo) and The Regolith and Ice Drill for Exploring New Terrain (TRIDENT)
 - Will fly to the south pole in late 2022 on an Intuitive Machines CLPS delivery

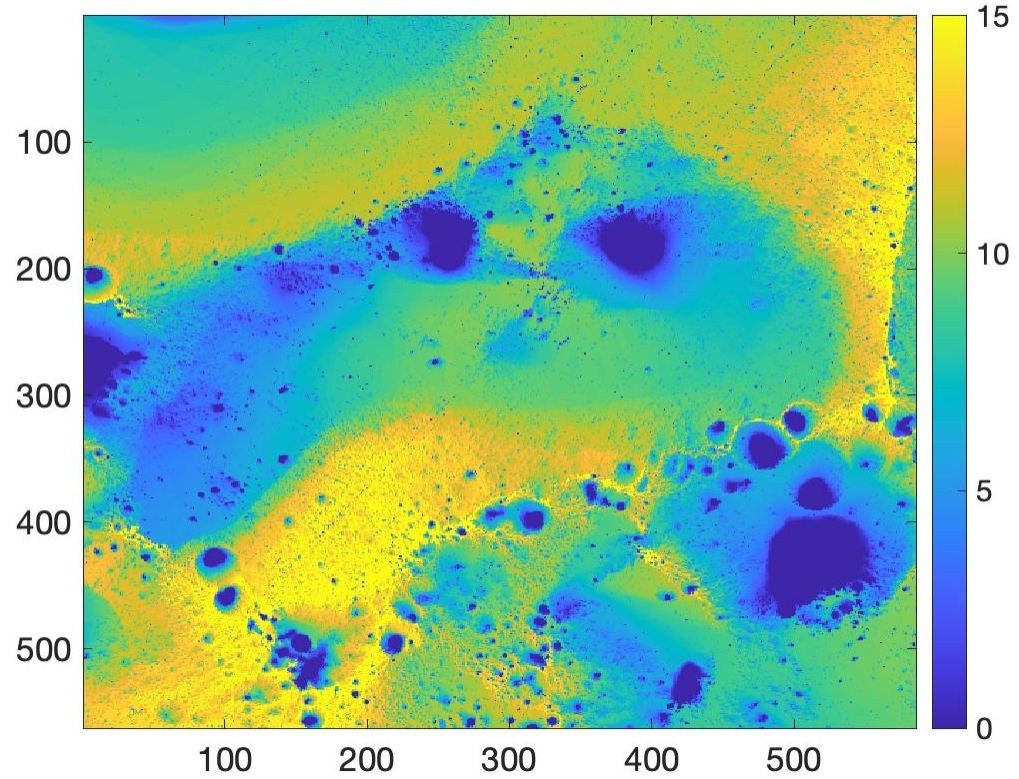
NASA is utilizing APL's expertise in precision illumination and thermal simulations to help select the best landing site for mission success



Longest period of continuous sunlight, starting January 2, 2023



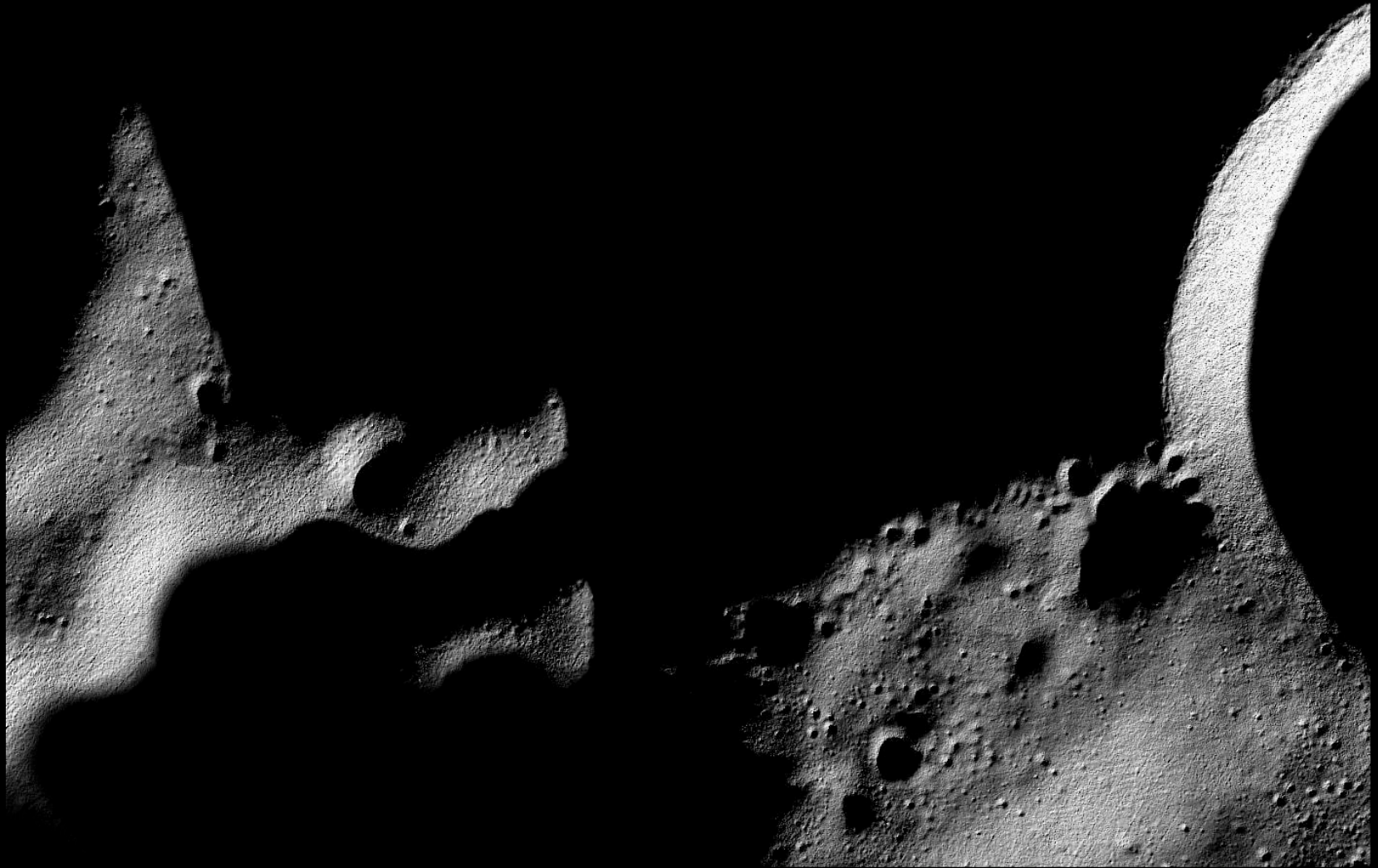
January 6, 2023



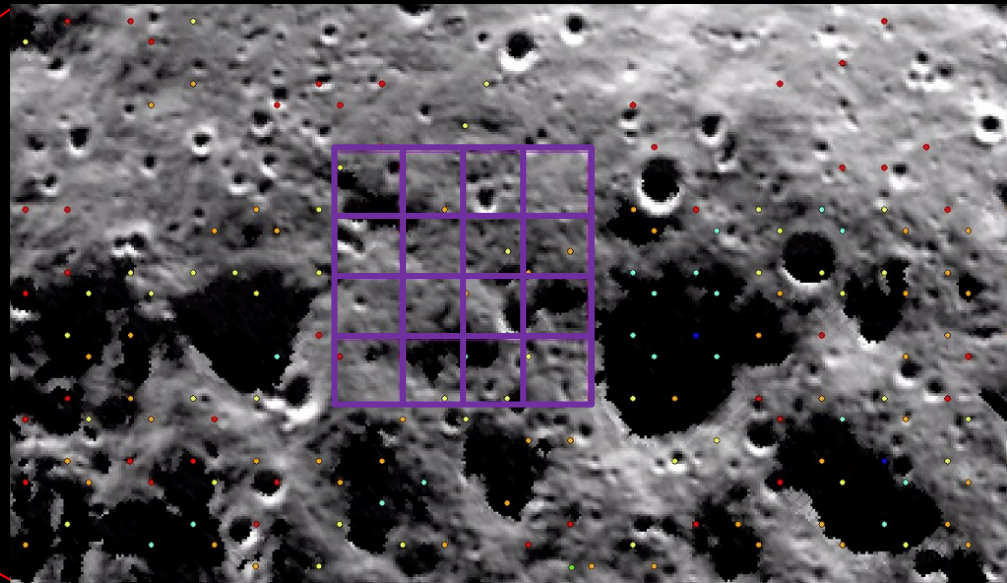
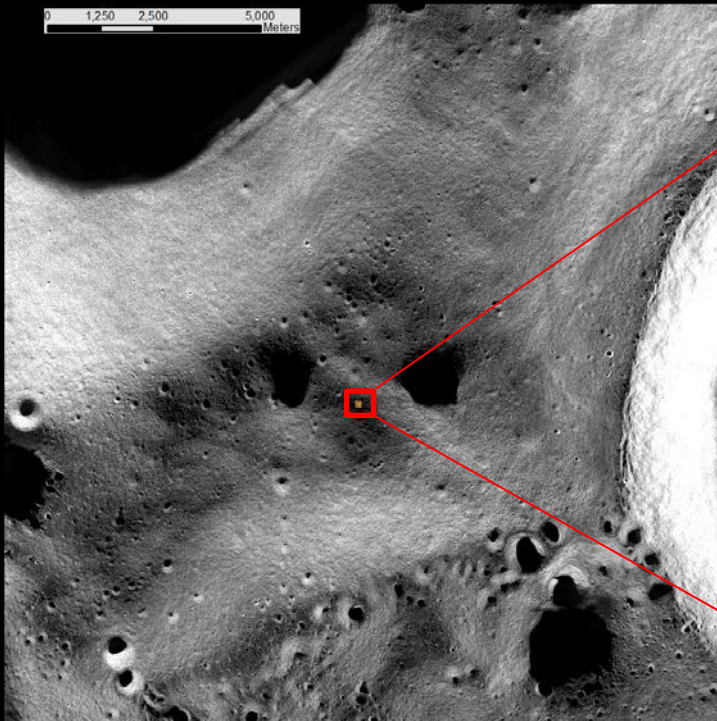
LOLA 25m DEM to 80S

Simulated Lighting Conditions – 25 m DEM

27-Dec-2022 00:00:00

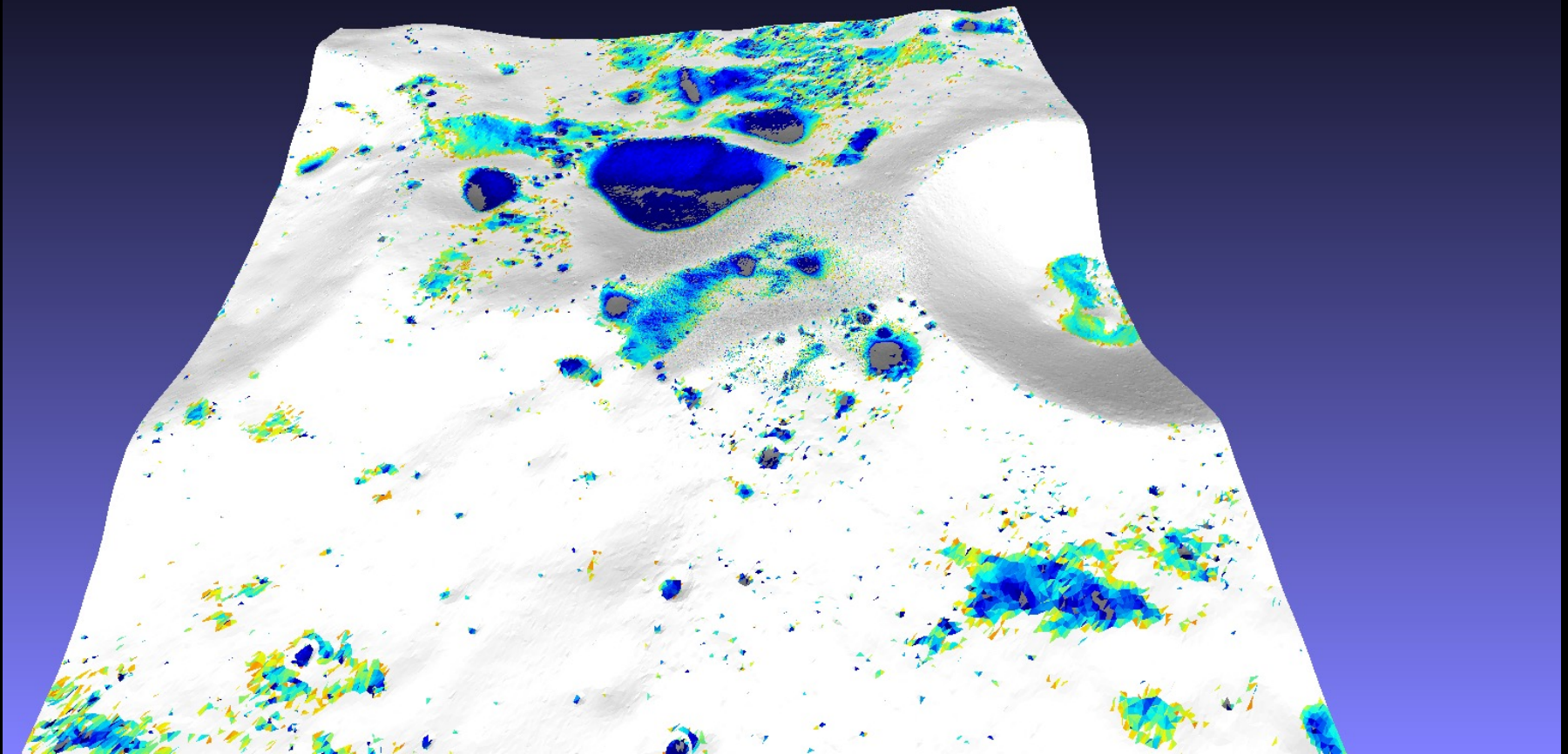


Valley Ridge Region

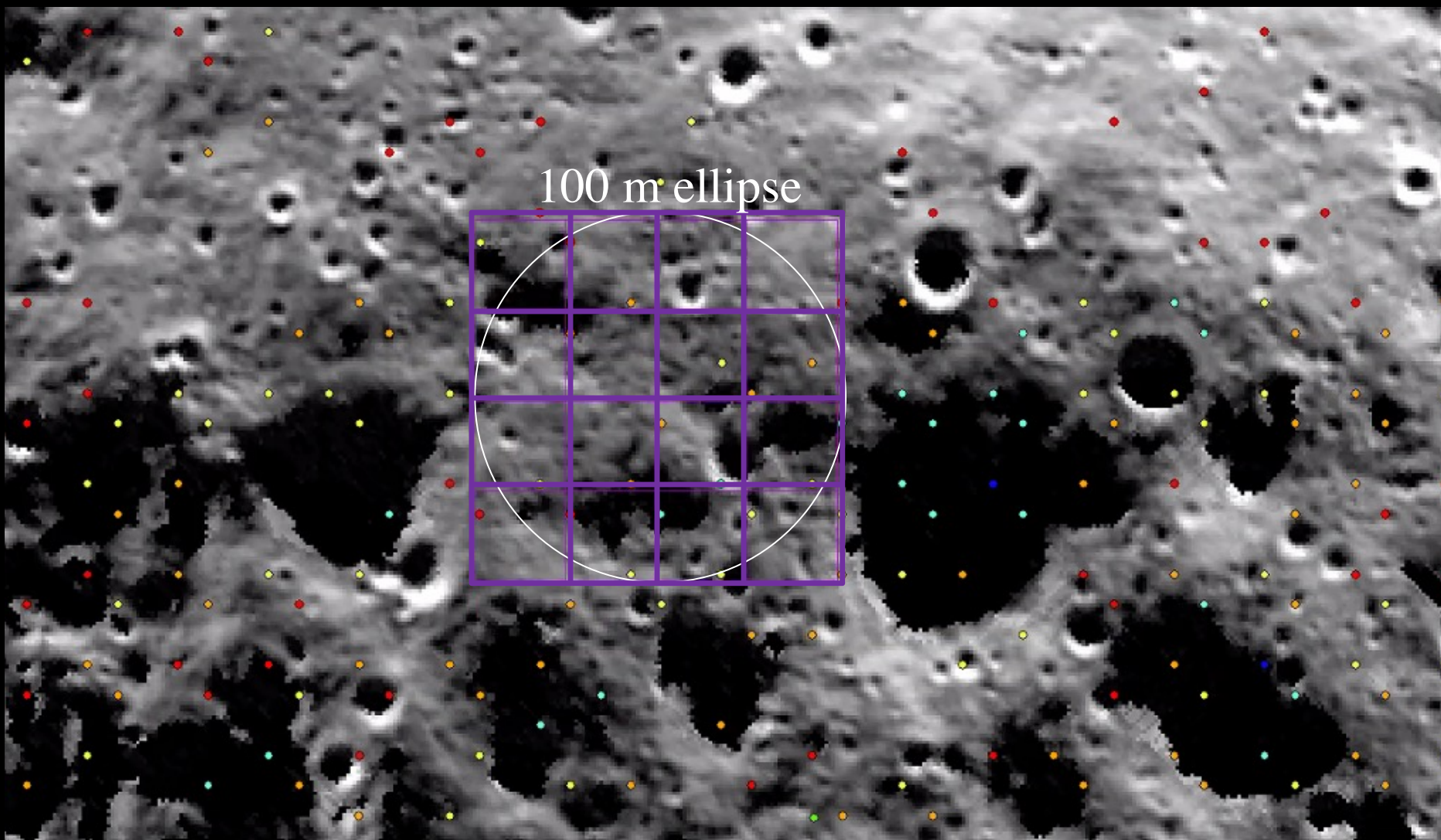


LOLA 25m DEM to 80S

Ice Stability

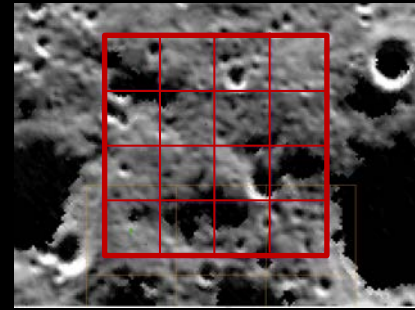


Valley Ridge Region, -89.6° , -125.4°



Valley Ridge Region

-89.6°, -125.4°



LOLA 25m DEM to 80S

1	5	9	13
2	6	10	14
3	7	11	15
4	8	12	16

Max total consecutive days sunlight

10	10	10	5
5	10	10	10
10	10	10	7
10	4	10	10

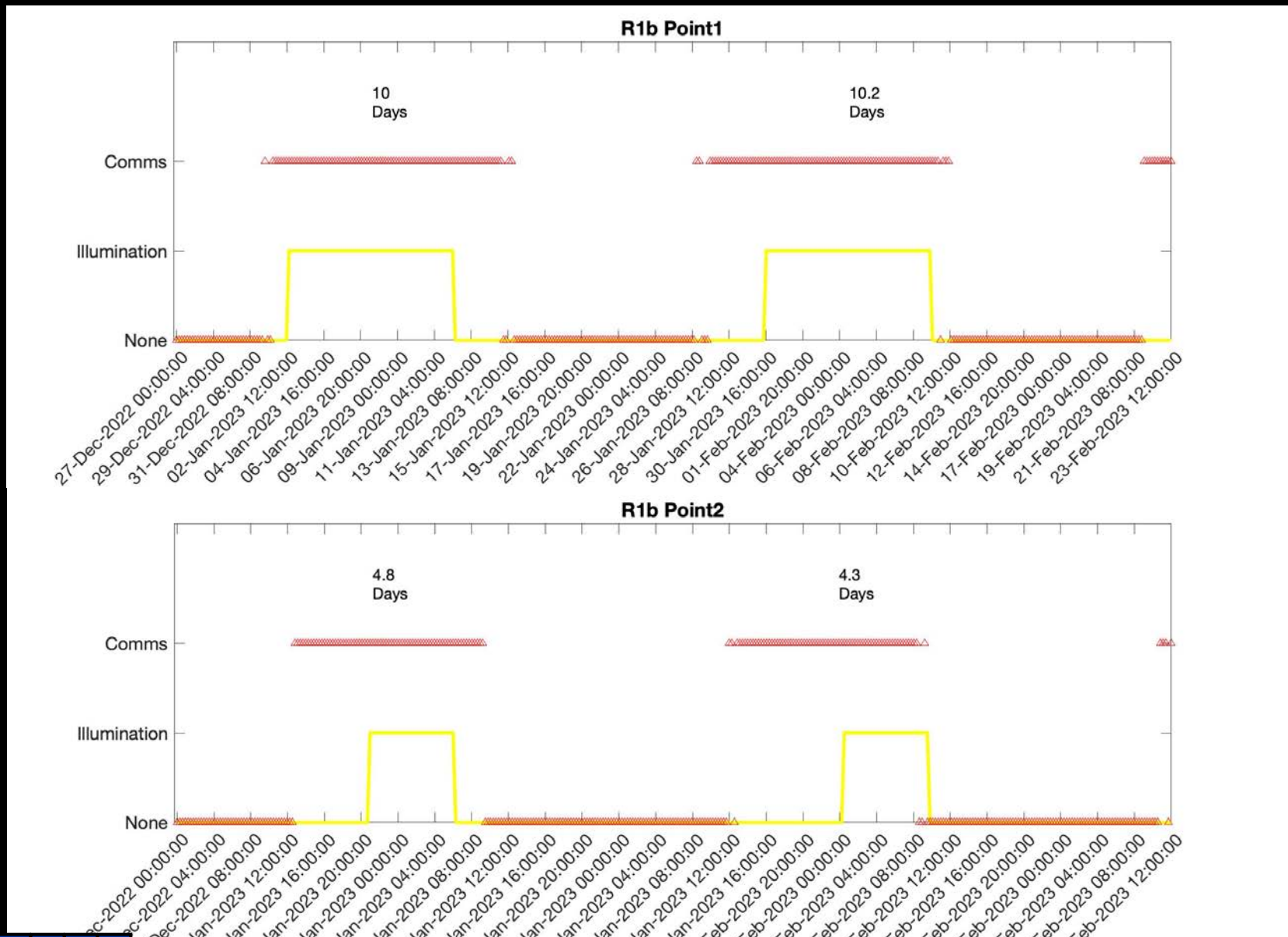
Jan

10	10	10	5
5	10	10	10
10	10	10	7
10	4	10	10

Feb

Point 1

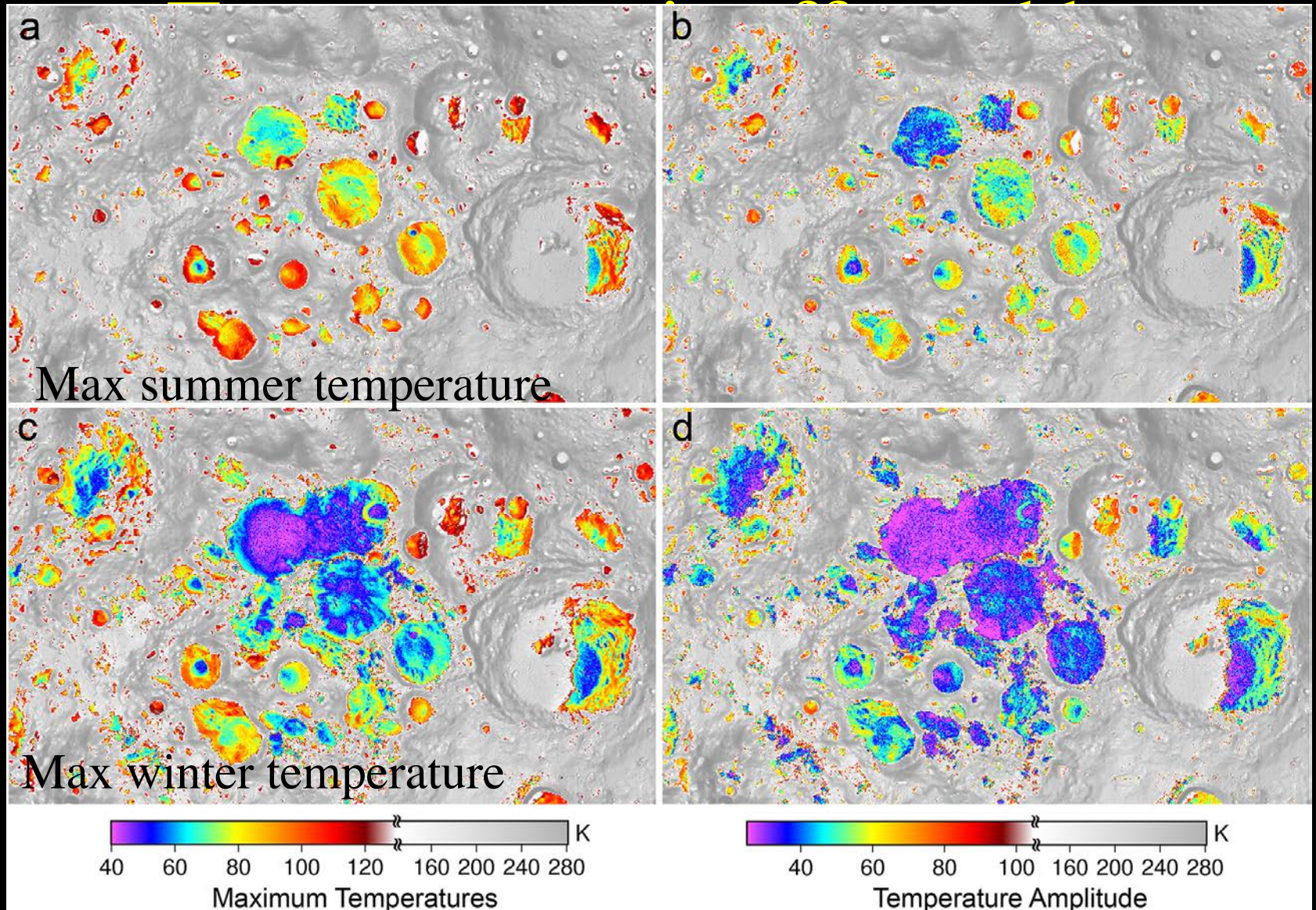
Point 2



1	5	9	13
2	6	10	14
3	7	11	15
4	8	12	16

LOLA 25m DEM to 80S

Ben Bussey / Angela Stickle JHU/APL



Seasonal temperatures from the Diviner instrument